



## Ground-Based Diurnal Measurements of $\text{NO}_2$ and $\text{NO}_3$ in Support of SAGE-III/ISS Validation

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# Objectives

**Objective 1:** Obtain 24 hour measurements of NO<sub>3</sub> and NO<sub>2</sub> from Table Mountain Facility in support of SAGE III/ISS validation.

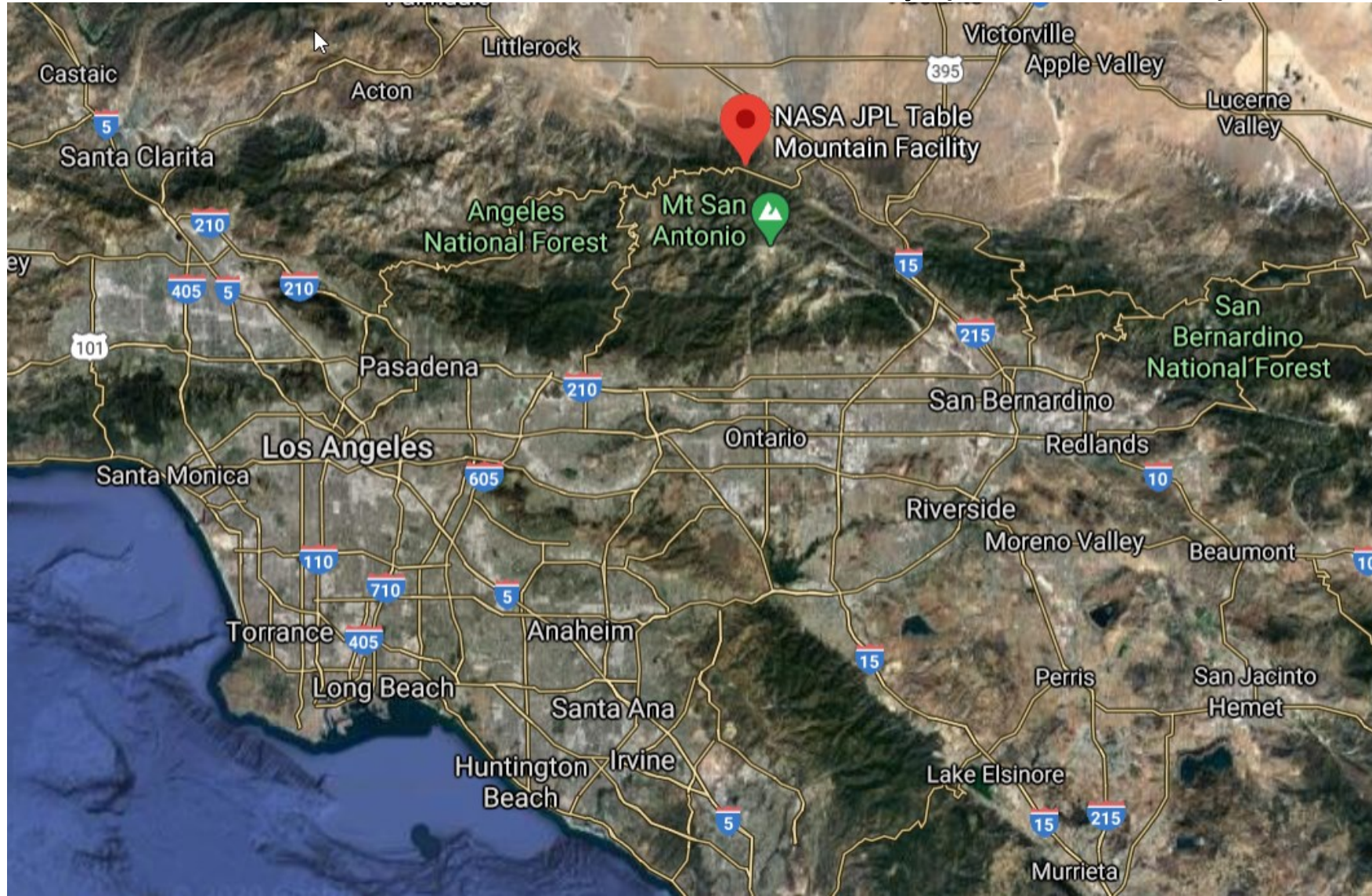
**Objective 2:** Analyze comparisons with coincident SAGE III/ISS measurements and work with the algorithm and science teams to interpret the comparisons and incorporate improvements into retrieval algorithms.

Roles		
Stanley P. Sander, JPL/Caltech	Data Acquisition / Analysis	Principal Investigator
Thomas J. Pongetti, JPL	Data Acquisition / Analysis	Co-Investigator
Yuk L. Yung, Caltech	1-D Modeling	Co-Investigator
King-Fai Li, UCR	1-D Modeling	Collaborator
Zhao-Cheng Zeng, UCLA/Caltech	Retrievals	Collaborator
Yangcheng Luo, Caltech	Radiative transfer / Retrievals	Collaborator



# Ground-Based Diurnal Measurements of $\text{NO}_2$ and $\text{NO}_3$

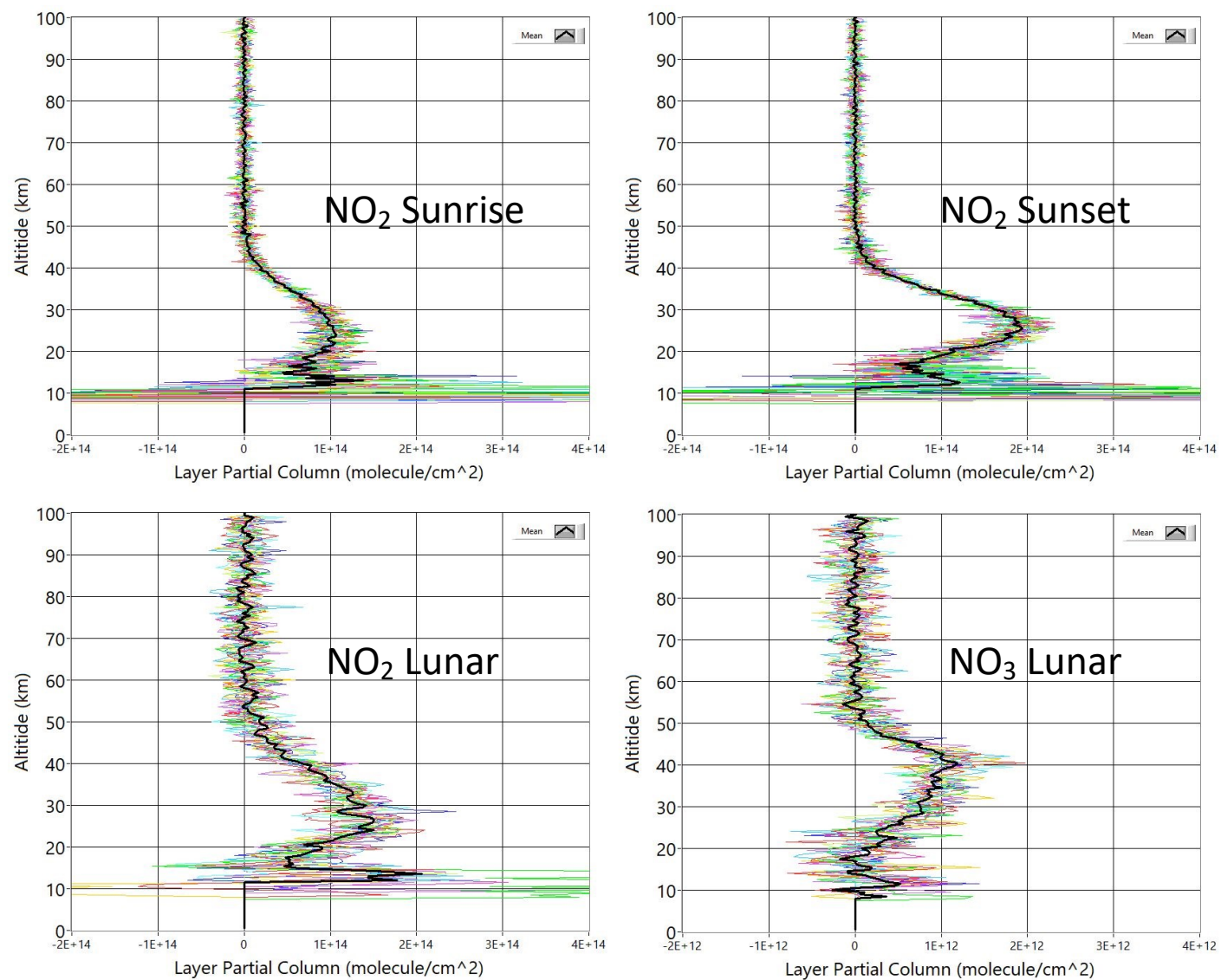
Direct Solar/Lunar and Sky light Spectroscopy  
at JPL Table Mountain Facility (2.3 km ASL)



# Issues Confronting Ground-Based Comparisons with SAGE III

- Lower limit of SAGE III vertical profiles is 12-15 km (due to large uncertainty only 17 km and above are used)
- Ground-based occultations extend from 2.3 km to TOA
- SAGE III solar occultations occur at sun rise/set
  - Diurnal Corrections described in Dubè et. al. have been used in this comparison
  - **We have implemented an additional measurement capability of zenith sky light to capture rise and set times which allows for a full 24 hour measurement of NO<sub>2</sub> (Radiative transfer model used to find zenith sky airmass factors)**
- Both SAGE III and ground-based lunar occultations occur over a range of local times

# MUGS Comparisons with SAGE-III/ISS

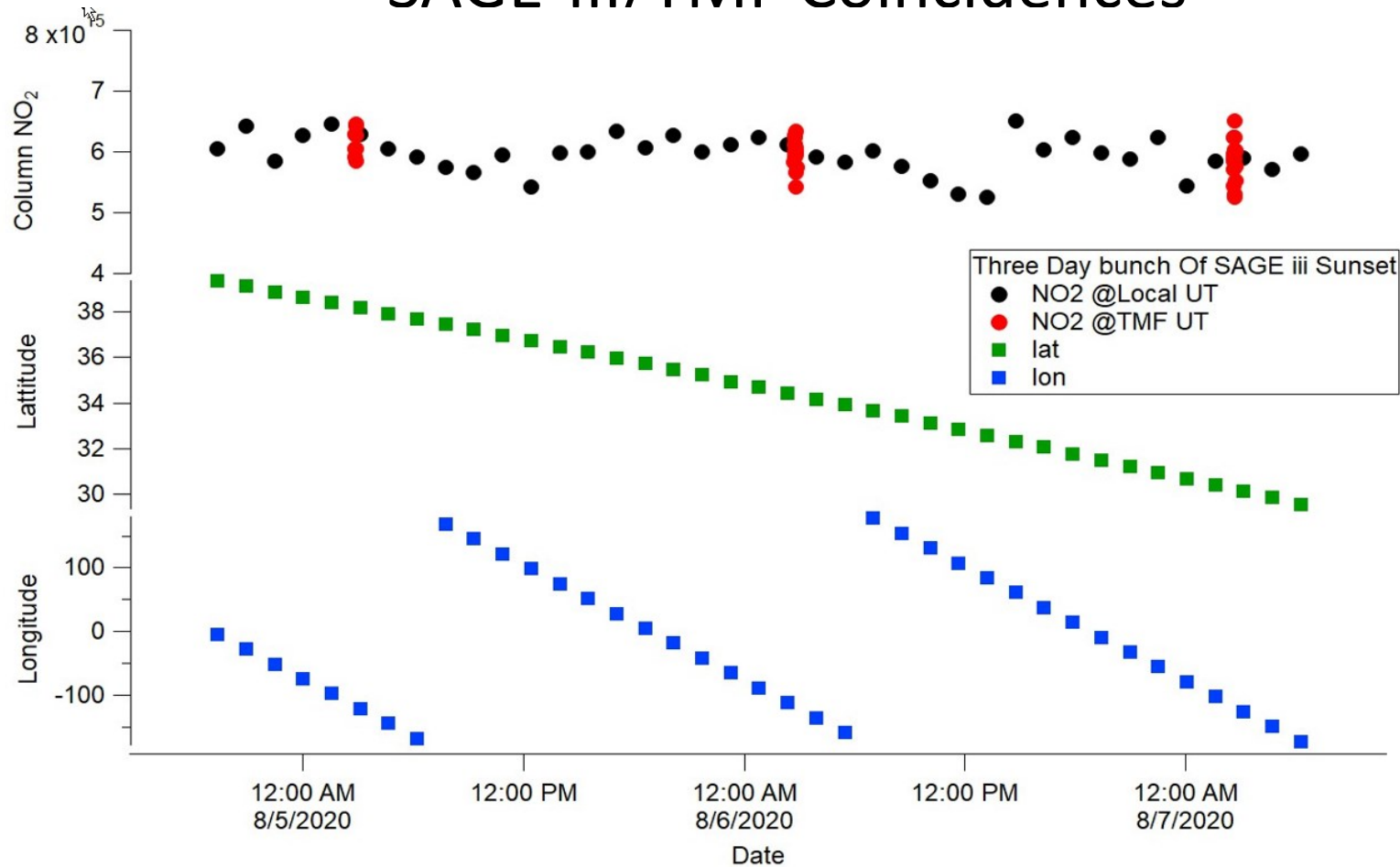


	Latitude Band	Number of Coincidence
NO <sub>2</sub> Sunrise	+/- 1°	333
NO <sub>2</sub> Sunset	+/- 1°	328
NO <sub>2</sub> Lunar	+/- 5°	366
NO <sub>3</sub> Lunar	+/- 5°	366

(SAGE III/ISS Profile Examples taken from single day from +/- 10°)

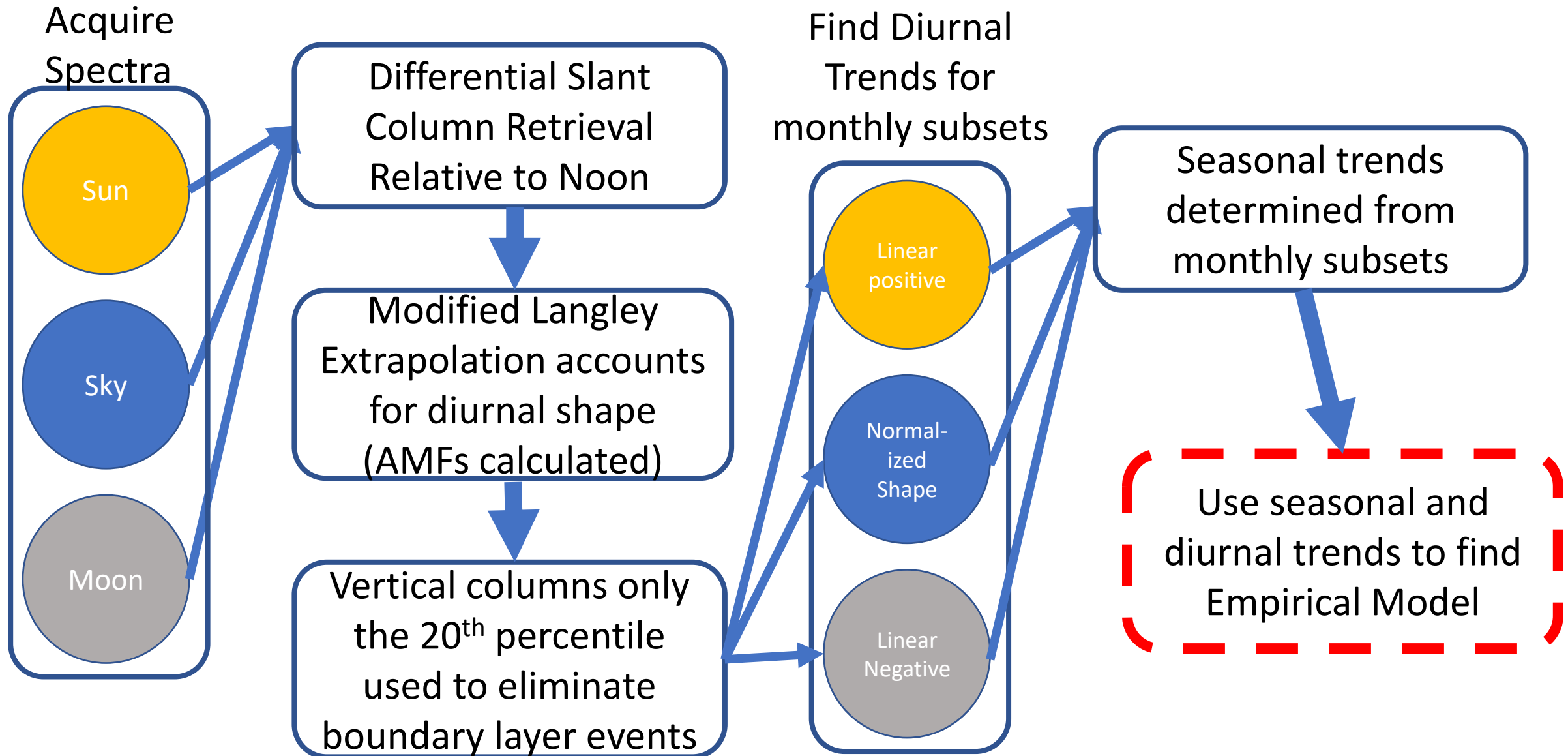


# SAGE-III/TMF Coincidences

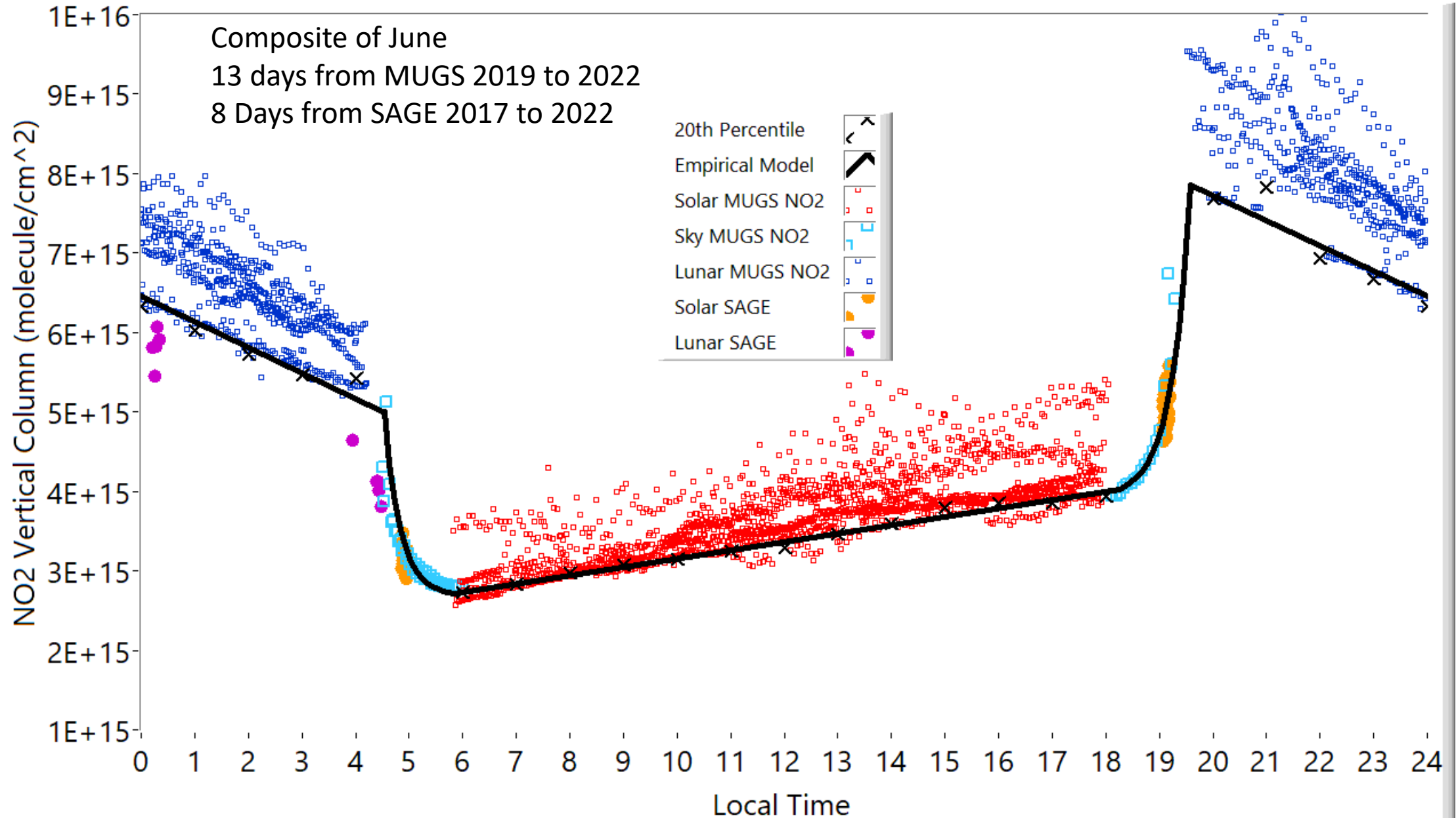


- Exact coincidences between SAGE-III and TMF are infrequent
- We consider coincidences within a global zone from TMF latitude(34.4 °N)
  - $\pm 1^\circ$  for solar and  $\pm 5^\circ$  latitude for lunar
- SAGE occultations at other longitudes are converted to TMF local time

# Creating an Empirical Model from the Measurements

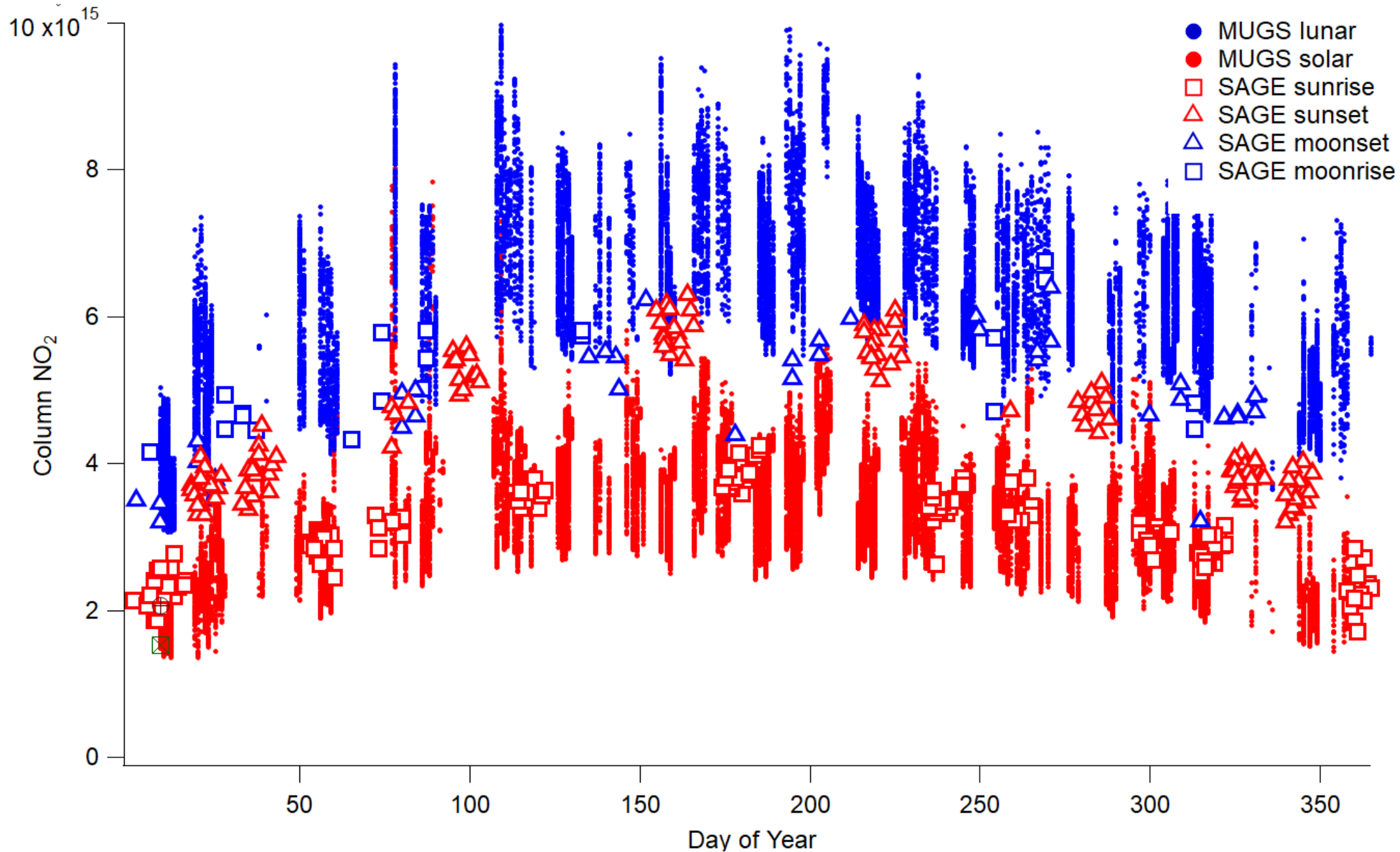


# Diurnal Variation of NO<sub>2</sub> Column: SAGE III and MUGS

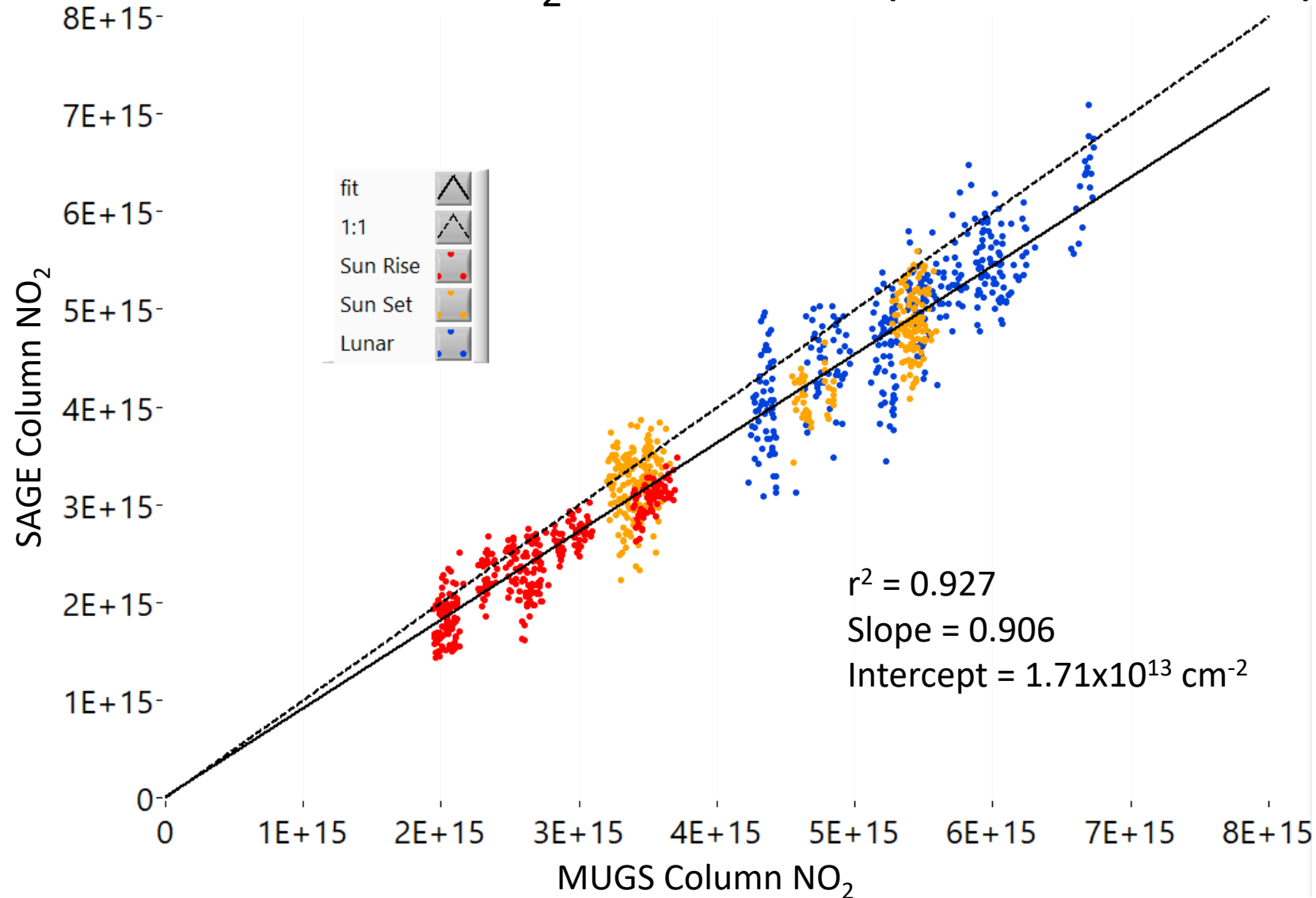




# SAGE-MUGS NO<sub>2</sub> Comparison: year-over-year

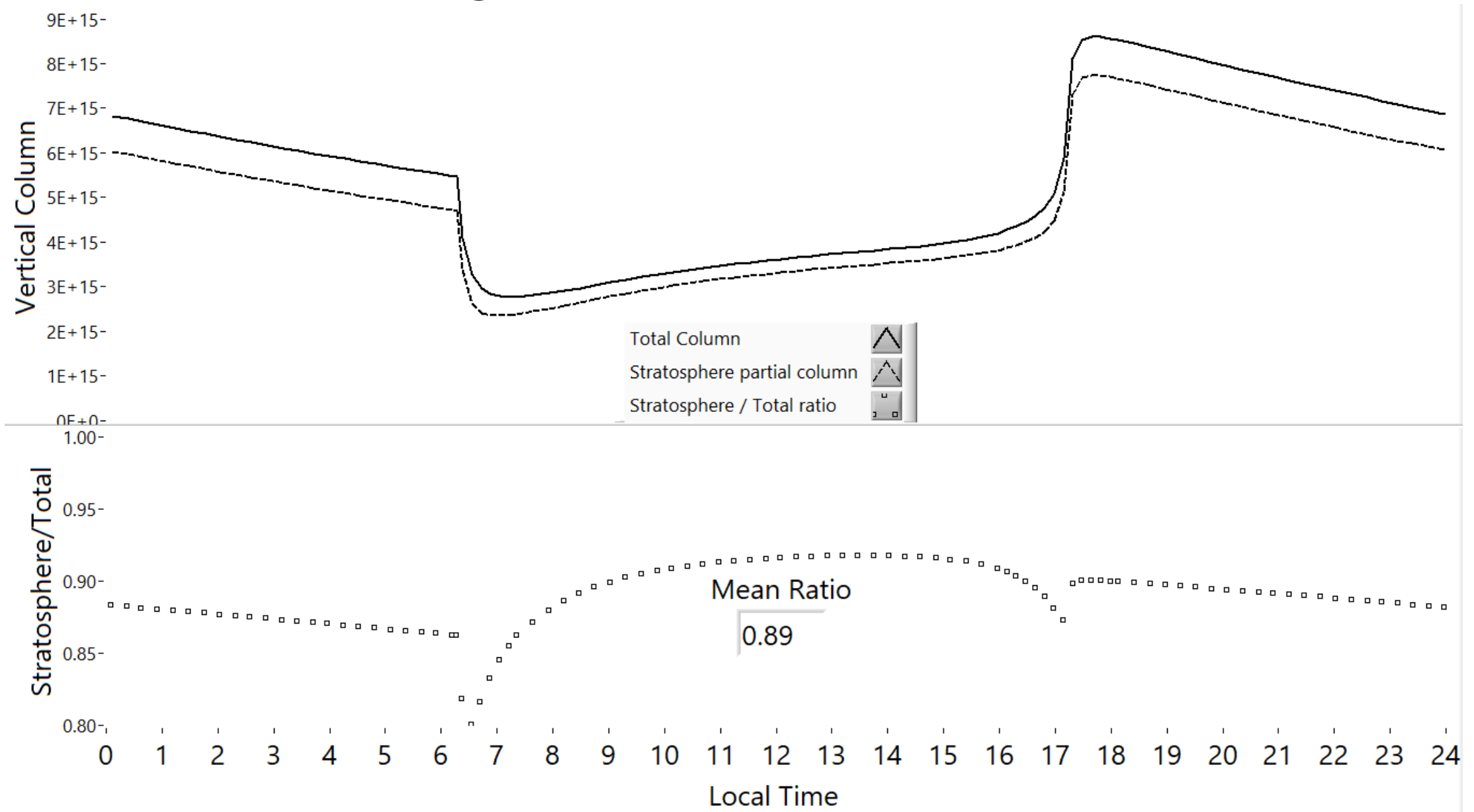


# SAGE-MUGS NO<sub>2</sub> Correlation (Solar and Lunar)



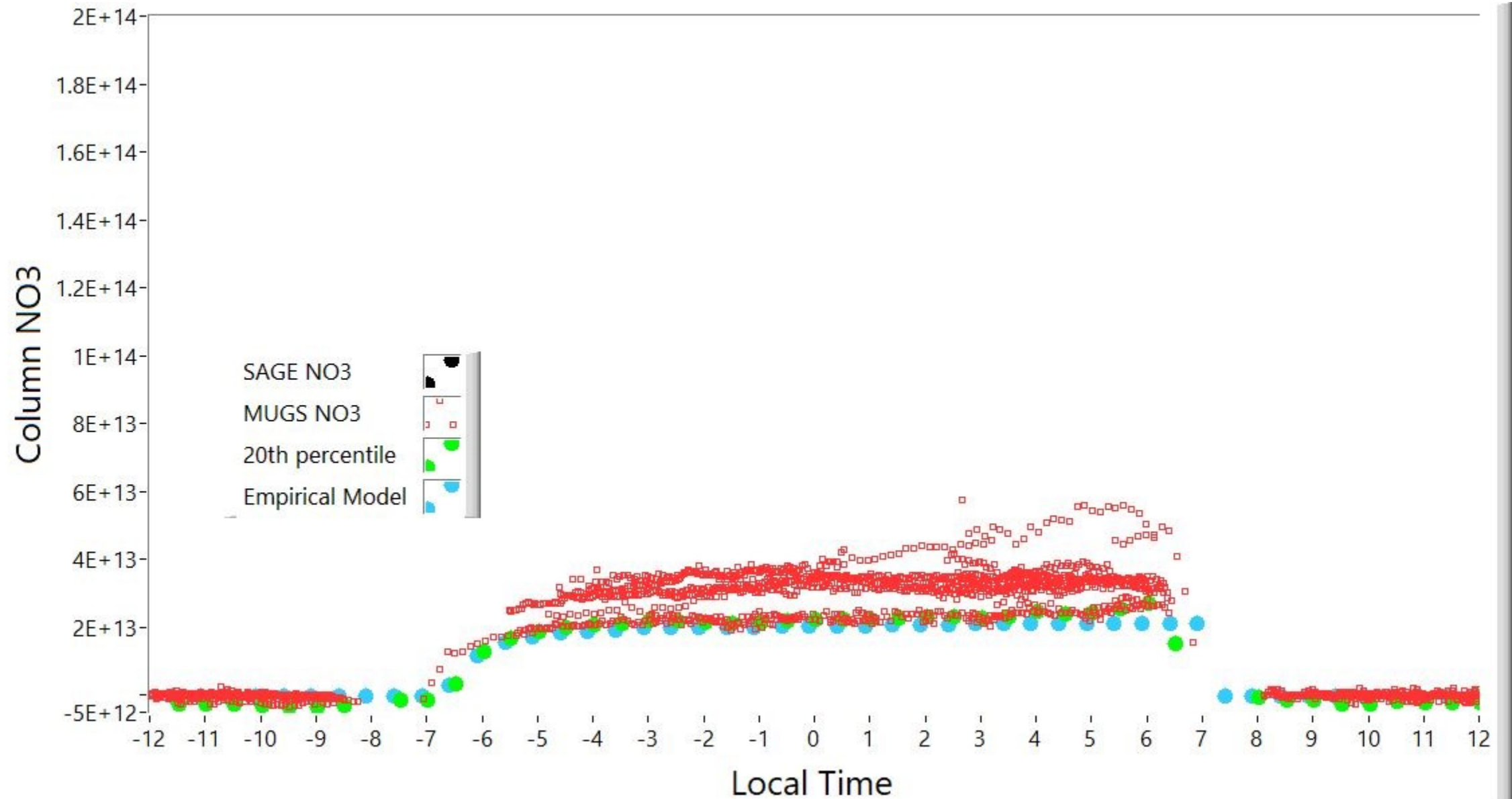
Scatter plot between SAGE-III partial columns ( $z > 17 \text{ km}$  and after diurnal correction)  
and MUGS total column empirical model

# Estimation of partial NO<sub>2</sub> column above 17 km Using the Photochemical Model

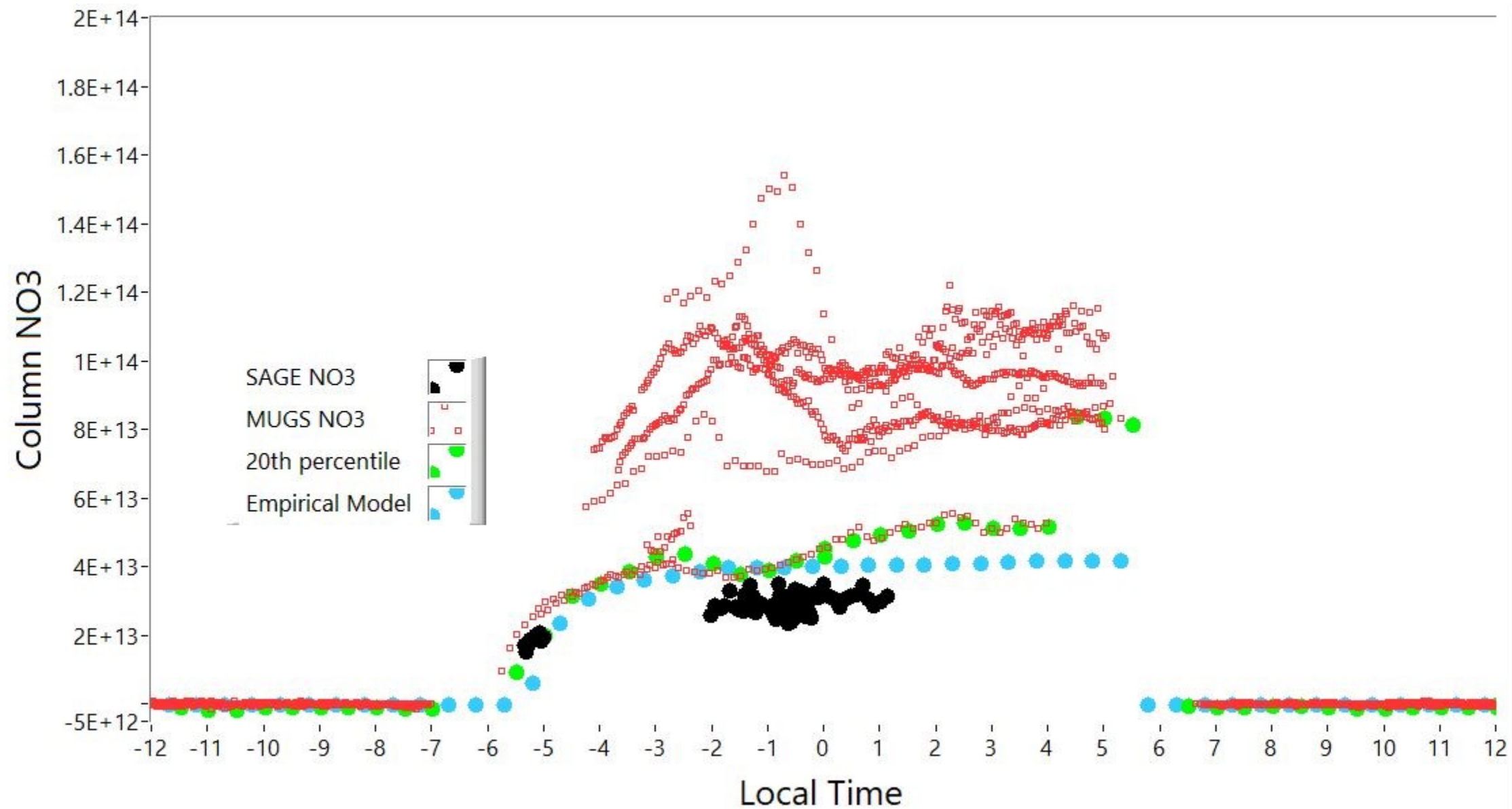




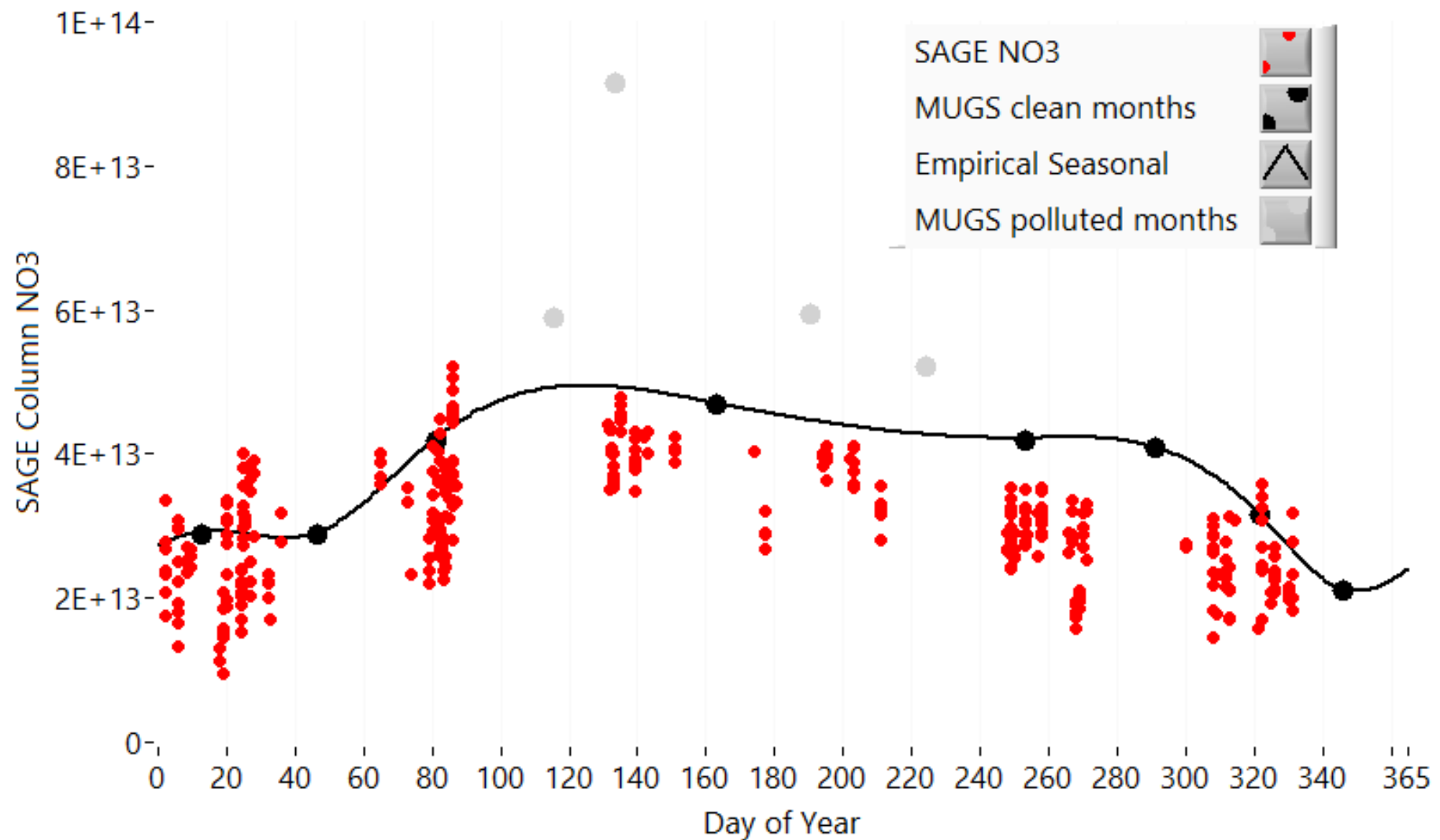
# MUGS NO<sub>3</sub> Compilation: December



# SAGE-MUGS NO<sub>3</sub> Comparison: September

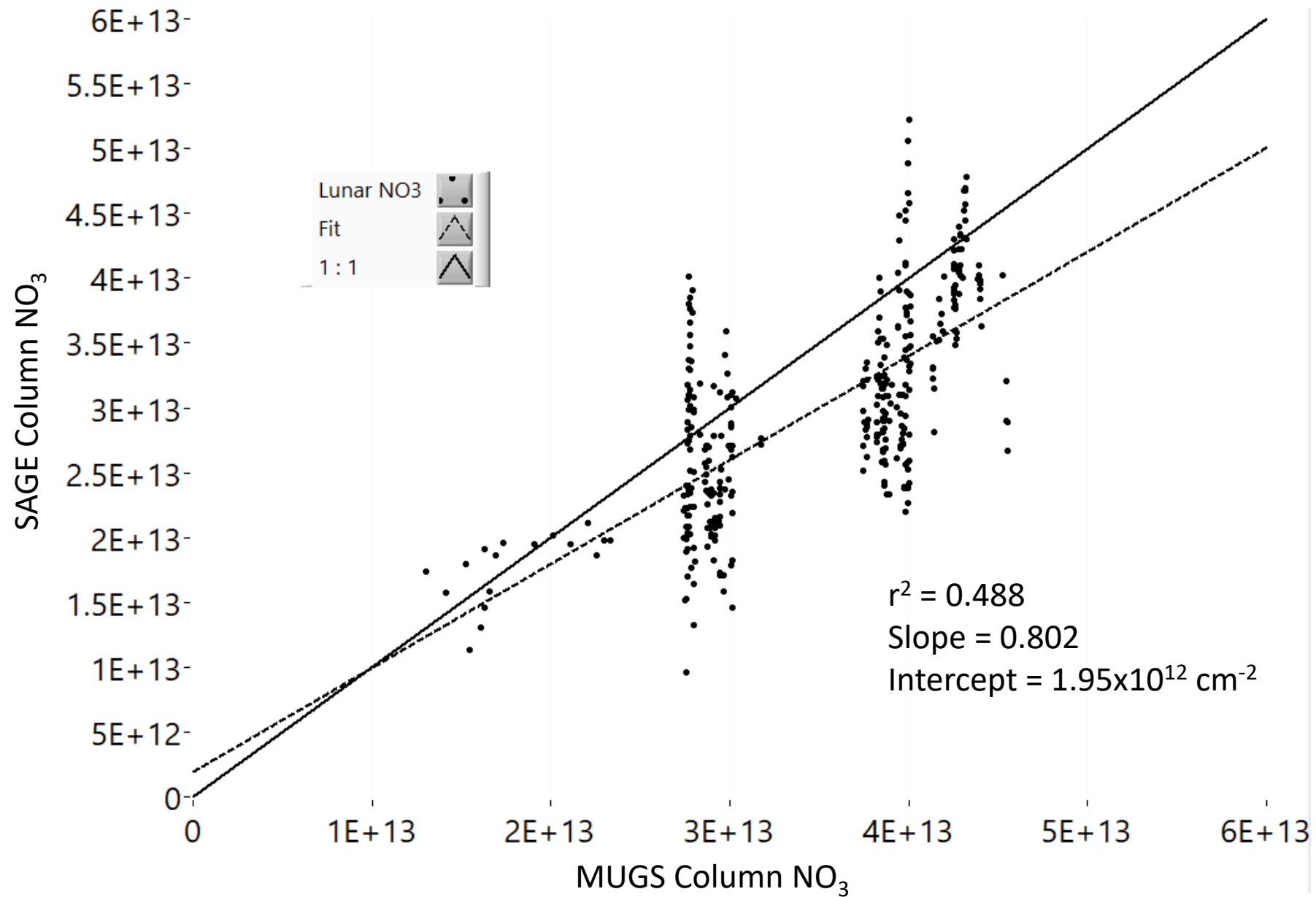


# SAGE-MUGS NO<sub>3</sub> Comparison: year-over-year





# SAGE-MUGS NO<sub>3</sub> Correlation



## Summary

- TMF total column measurements
  - Effects of PBL pollution, diurnal trends, and seasonal variation have been corrected using the entirety of the MUGS data set
- SAGE-III profile measurements are converted to partial column
  - Diurnally corrected data from Latitude bands surrounding TMF collapsed to local time currently provide 1393 individual profiles to compare
- NO<sub>2</sub> shows excellent correlation between the two data sets with SAGE-III being systematically lower by about 10%. This is almost perfectly accounted for by the photochemical model showing the same amount is below 17 km.
- NO<sub>3</sub> has decent correlation to between the two data sets with SAGE-III being systematically lower by about 20%. Further focus on this to come.
- *Future work* will extending the ground-based climatologies and diurnal trends, correcting for FT partial column using more comprehensive model analysis, studying the radiative transfer model to improve ground based measurements of sunrise and sunset. Collection of more direct coincidence measurements.